

# Physics

## Today's Agenda

- Reference frames and relative motion.
- Uniform Circular Motion

## Inertial Reference Frames:

- A **Reference Frame** is the place you measure from.
  - } It's where you nail down your  $(x,y,z)$  axes!
- An Inertial Reference Frame (**IRF**) is one that is **not accelerating** with respect to the “fixed stars”.
  - } We will consider only **IRF**'s in this course.
  - } The earth is a pretty good **IRF** even though it is accelerating (rotating) w.r.t. the stars slightly.
- Valid **IRF**'s can have fixed velocities w.r.t each-other.
  - } More about this later when we discuss forces.
  - } For now, just remember that we can make measurements from different vantage points.

## Relative Motion

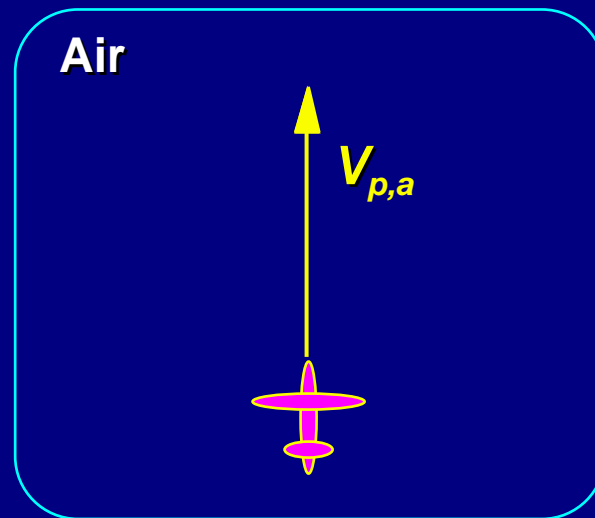
- Consider a problem with **two** distinct IRF's:
  - } **An airplane flying on a windy day.**

A pilot wants to fly from Austin to Dallas. Having asked a friendly student, she knows that Dallas is 120 miles due north of Austin. She takes off from the Airport at noon. Her plane has a compass and an air-speed indicator to help her navigate.

- } The compass allows her to keep the nose of the plane pointing north.
- } The air-speed indicator tells her that she is traveling at 120 miles per hour **with respect to the air.**

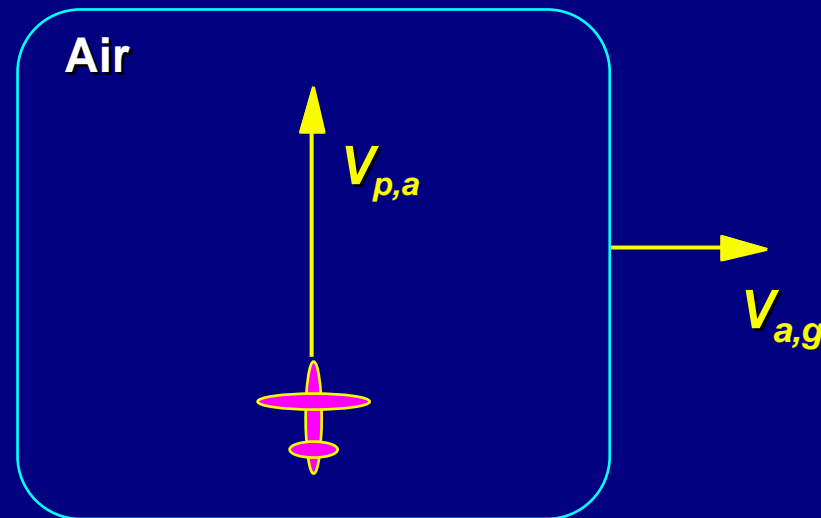
## Relative Motion...

- The plane is moving north in the IRF attached to the air:  
}  $V_{p,a}$  is the velocity of the plane w.r.t the air.



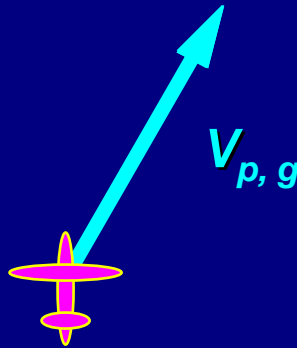
## Relative Motion...

- But suppose the air is moving *east* in the IRF attached to the ground.
  - }  $V_{a,g}$  is the velocity of the air w.r.t the ground (i.e. *wind*).



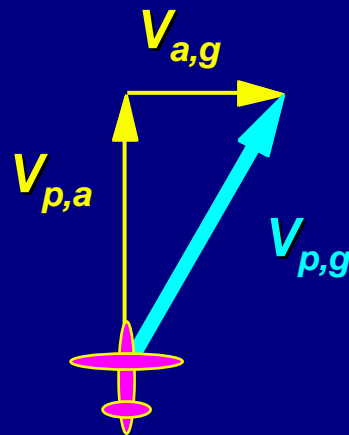
## Relative Motion...

- What is the velocity of the plane in an IRF attached to the ground?  
}  $V_{p, g}$  is the velocity of the plane w.r.t the ground.

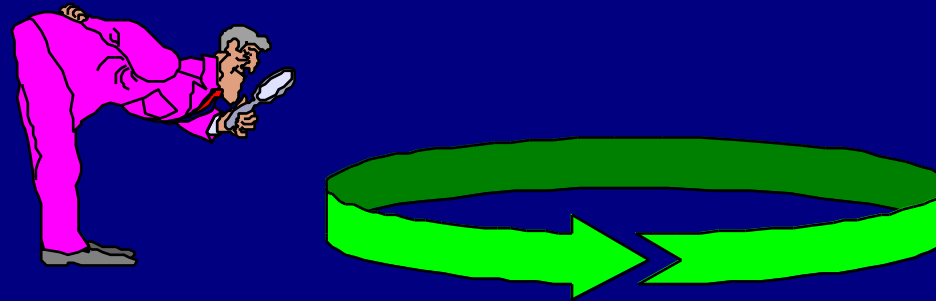


## Relative Motion...

$V_{p,g} = V_{p,a} + V_{a,g}$  Is a vector equation relating the airplanes velocity in different reference frames.



# Uniform Circular Motion

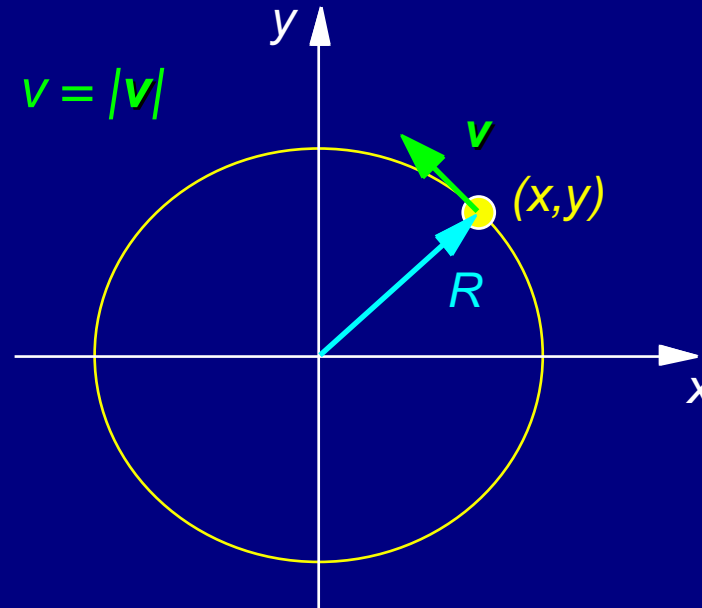


- What does it mean ?
- How do we describe it ?
- What can we learn about it ?



## What is UCM?

- Motion in a circle with:
  - } Constant Radius  $R$
  - } Constant Speed  $v = |\mathbf{v}|$



## How can we describe UCM?

- In general, one co-ordinate system is as good as any other:

- } Cartesian:

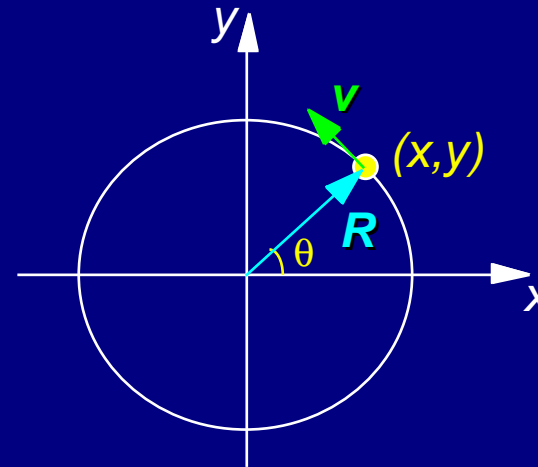
- »  $(x,y)$  [position]

- »  $(v_x, v_y)$  [velocity]

- } Polar:

- »  $(R,\theta)$  [position]

- »  $(v_R, \omega)$  [velocity]



- In UCM:

- }  $R$  is constant (hence  $v_R = 0$ ).

- }  $\omega$  (angular velocity) is constant.

- } Polar co-ordinates are a natural way to describe UCM!

## Polar Coordinates:

- The arc length  $s$  (distance along the circumference) is related to the angle in a simple way:

$s = R\theta$ , where  $\theta$  is the *angular displacement*.

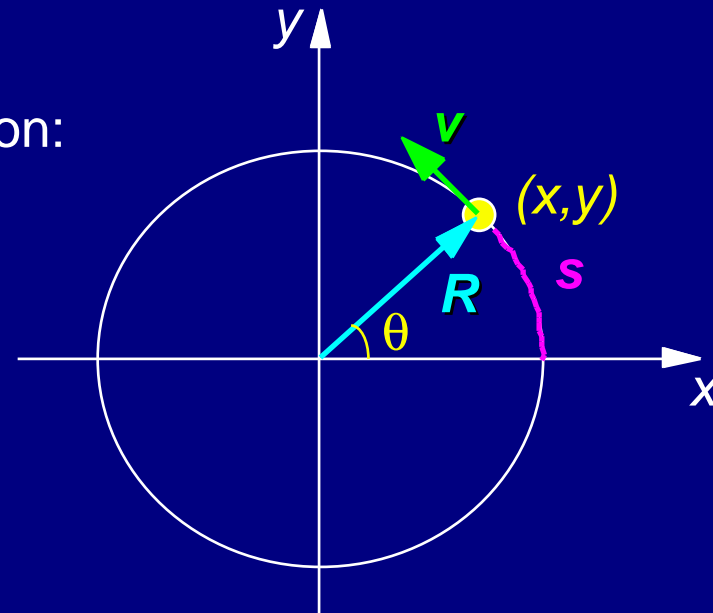
} units of  $\theta$  are called *radians*.

- For one complete revolution:

$$2\pi R = R\theta_c$$

}  $\theta_c = 2\pi$

$\theta$  has *period*  $2\pi$ .

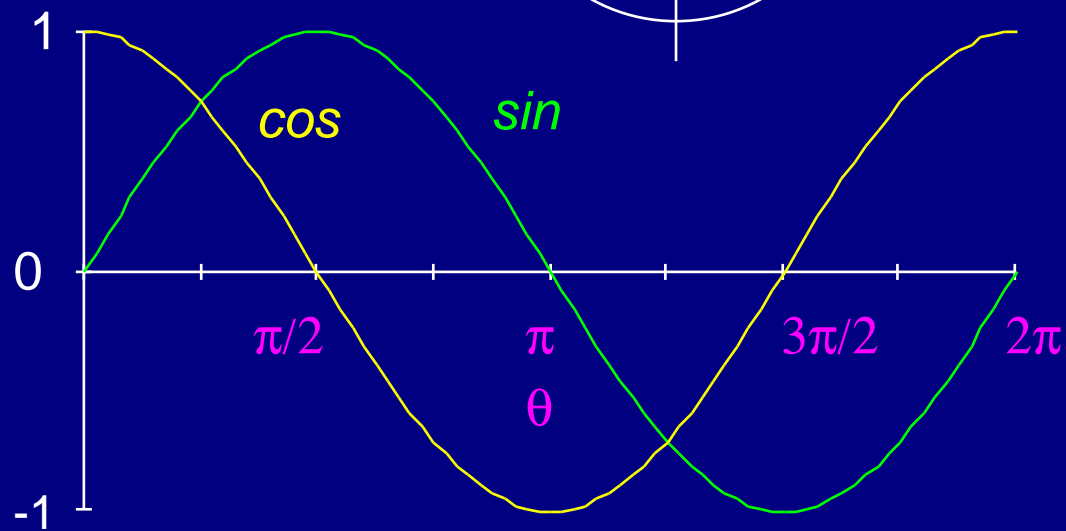
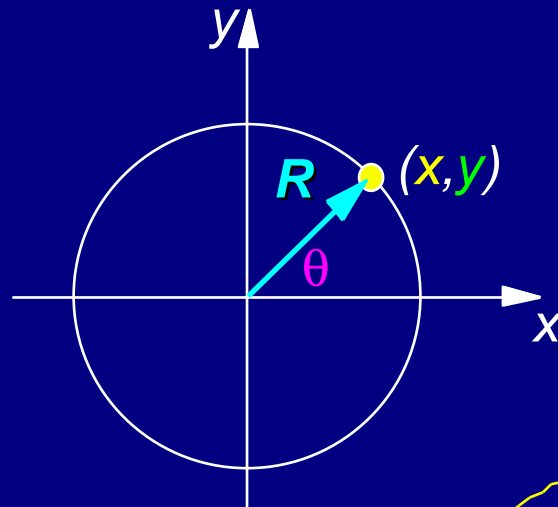


} 1 revolution =  $2\pi$  radians.

## Polar Coordinates...

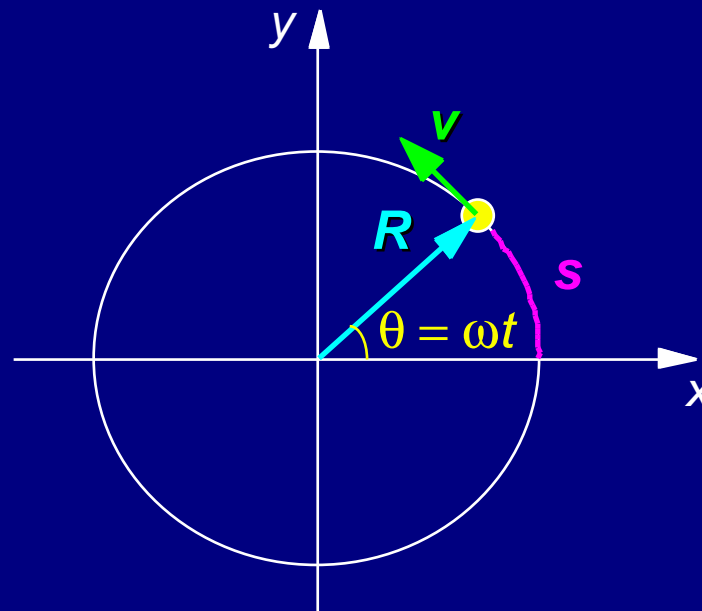
$$x = R \cos(\theta)$$

$$y = R \sin(\theta)$$



## Polar Coordinates...

- In cartesian co-ordinates we say velocity  $dx/dt = v$ .  
}  $x = vt$
- In polar coordinates, angular velocity  $d\theta/dt = \omega$ .  
}  $\theta = \omega t$
- Displacement  $s = vt$ .  
but  $s = R\theta = R\omega t$ .  
} so  $v = R\omega$



## Recap:

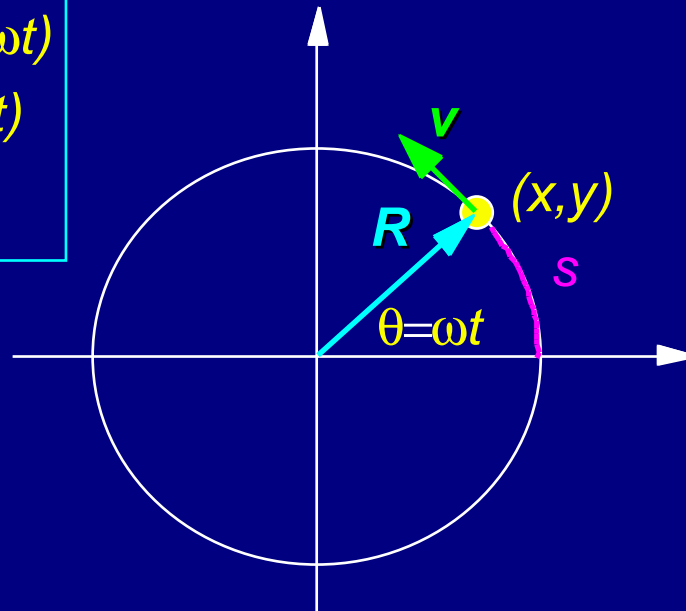
$$\begin{aligned}x &= R \cos(\theta) = R \cos(\omega t) \\y &= R \sin(\theta) = R \sin(\omega t) \\\theta &= \tan^{-1}(y/x)\end{aligned}$$

$$\theta = \omega t$$

$$s = vt$$

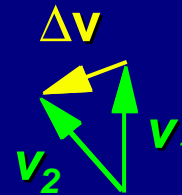
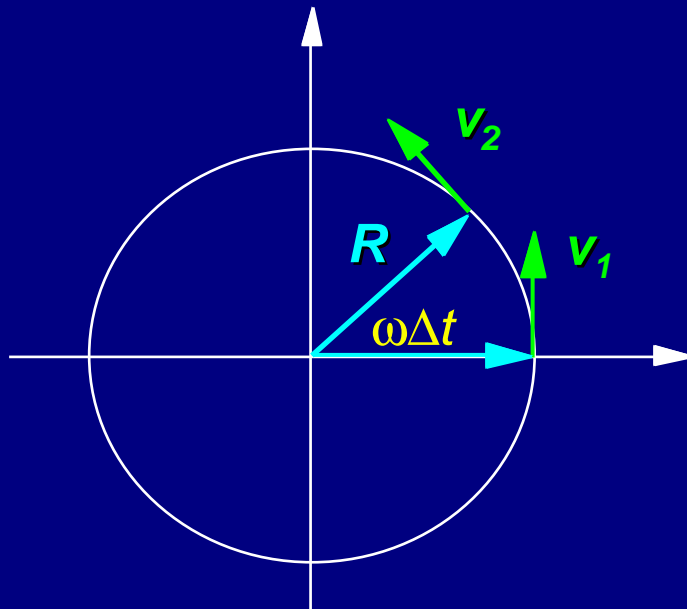
$$s = R\theta = R\omega t$$

$$v = R\omega$$



## Acceleration in UCM:

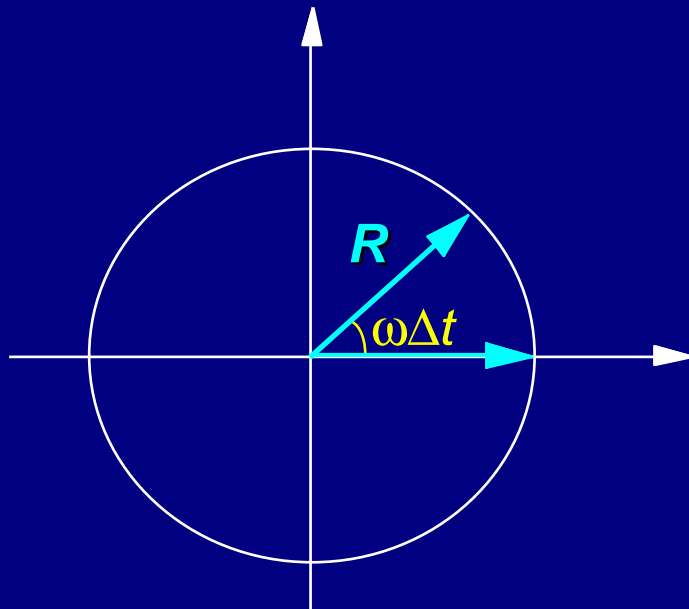
- Even though the *speed* is constant, *velocity* is *not* constant since the direction is changing.
  - } Consider average acceleration in time  $\Delta t \Rightarrow \Delta \mathbf{v} / \Delta t$



## Acceleration in UCM:

- Even though the **speed** is constant, **velocity** is **not** constant since the direction is changing.

} Consider average acceleration in time  $\Delta t$   $\Rightarrow \Delta \mathbf{v} / \Delta t$

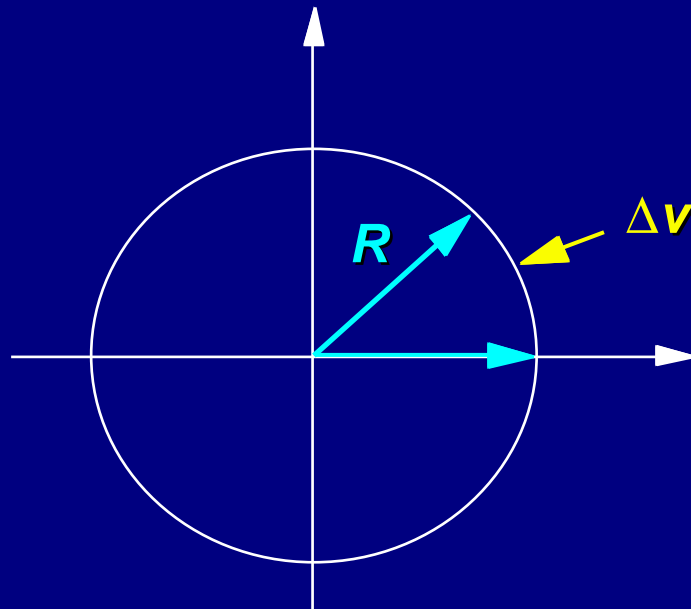


Move  $\Delta \mathbf{v}$  to average time of the interval  $\Delta t$ .



## Acceleration in UCM:

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  - } Consider average acceleration in time  $\Delta t \Rightarrow \Delta \mathbf{v} / \Delta t$

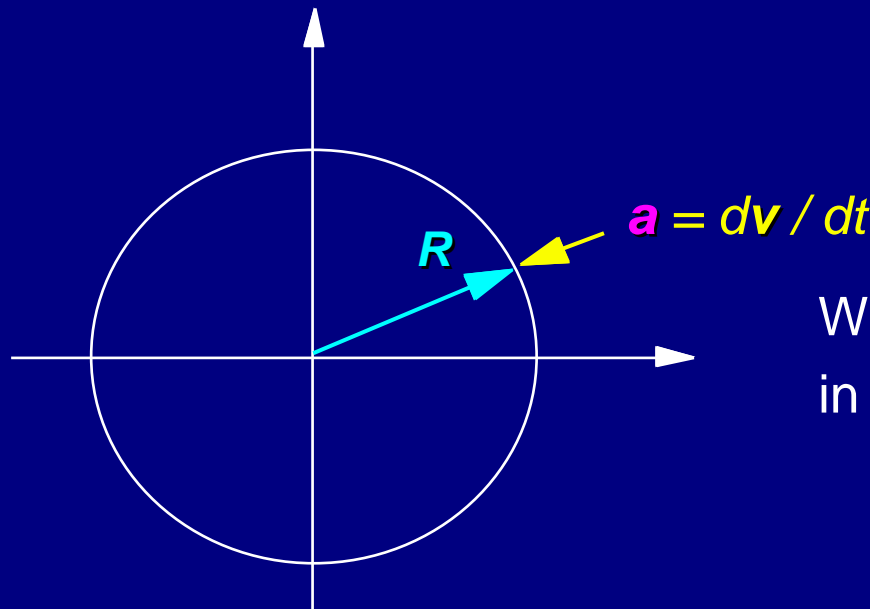


seems like  $\Delta \mathbf{v}$  (hence  $\Delta \mathbf{v} / \Delta t$ )  
points toward the origin !

## Acceleration in UCM:

- Even though the **speed** is constant, **velocity** is **not** constant since the direction is changing.

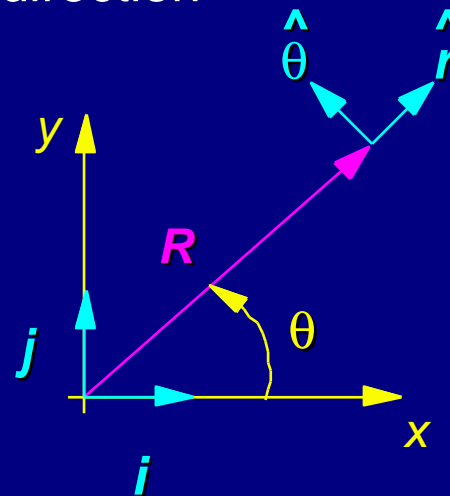
} As we shrink  $\Delta t$ ,  $\Delta \mathbf{v} / \Delta t \Rightarrow d\mathbf{v} / dt = \mathbf{a}$



We see that  $\mathbf{a}$  points in the  $-R$  direction.

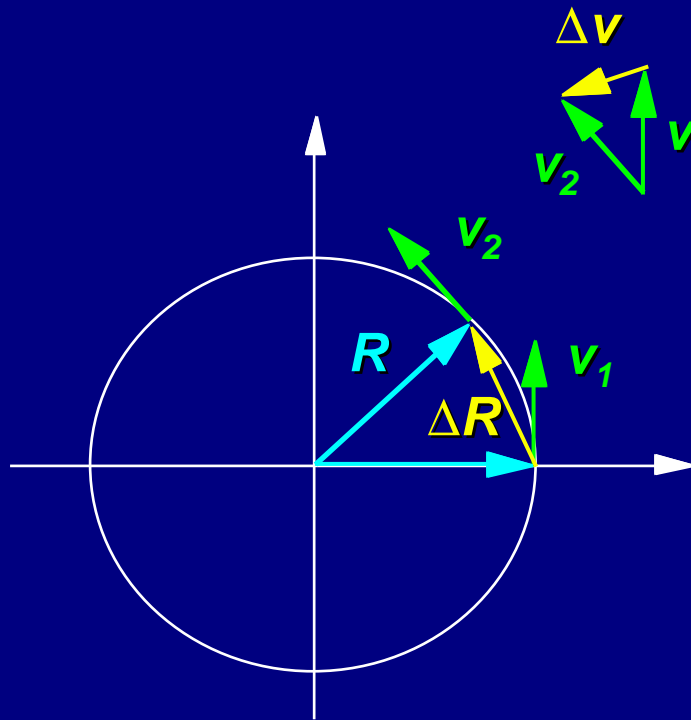
## Aside: Polar Unit Vectors

- We are familiar with the cartesian unit vectors:  $\hat{i}$   $\hat{j}$   $\hat{k}$
- Now introduce “polar unit-vectors”  $\hat{r}$  and  $\hat{\theta}$  :
  - }  $\hat{r}$  points in radial direction
  - }  $\hat{\theta}$  points in tangential (ccw) direction



## Centripetal Acceleration

- Must be accelerating if direction is changing:  
} Centripetal Acceleration !!



Similar triangles:  $\frac{\Delta v}{v} = \frac{\Delta R}{R}$

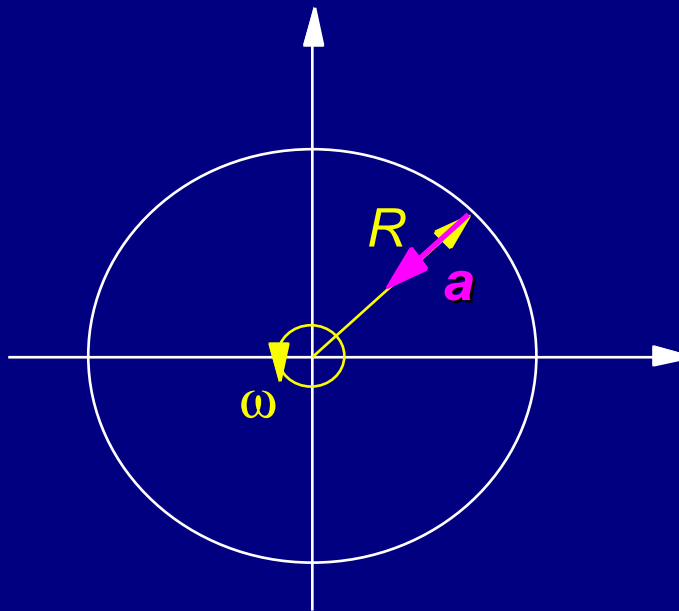
But  $\Delta R = v\Delta t$  for small  $\Delta t$

So:  $\frac{\Delta v}{v} = \frac{v\Delta t}{R} \Rightarrow \frac{\Delta v}{\Delta t} = \frac{v^2}{R}$

$$a = \frac{v^2}{R}$$

# Centripetal Acceleration

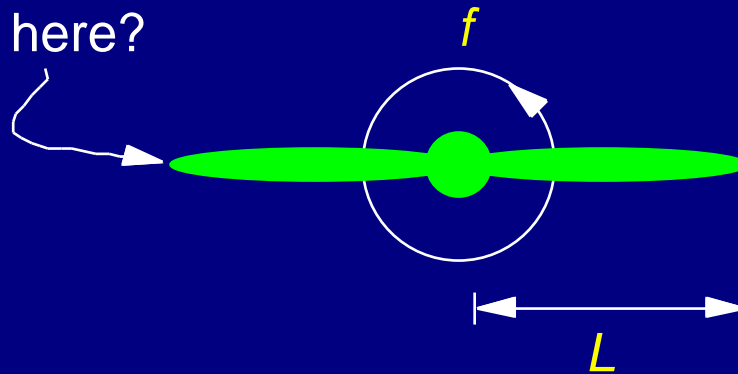
- UCM results in acceleration:
  - } Magnitude:  $a = v^2 / R$  ( $= \omega^2 R$  since  $v = R\omega$ )
  - } Direction:  $-\hat{r}$  (toward center of circle)



## Example: Propeller Tip

- The propeller on a stunt plane spins with frequency  $f = 3500 \text{ rpm}$ . The length of each propeller blade is  $L = 80 \text{ cm}$ . What centripetal acceleration does a point at the tip of a propeller blade feel?

what is  $a$  here?



## Example:

- First calculate the angular velocity of the propeller:

$$\} \quad 1 \text{ rpm} = 1 \frac{\text{rot}}{\text{min}} \times \frac{1 \text{ min}}{60 \text{ s}} \times 2\pi \frac{\text{rad}}{\text{rot}} = 0.105 \frac{\text{rad}}{\text{s}} \equiv 0.105 \text{ s}^{-1}$$

$$\} \quad \text{so } 3500 \text{ rpm means } \omega = 367 \text{ s}^{-1}$$

- Now calculate the acceleration.

$$\} \quad \mathbf{a} = \omega^2 R = (367 \text{ s}^{-1})^2 \times (0.8 \text{ m}) = 1.1 \times 10^5 \text{ m/s}^2 \\ = 11,000 \text{ g}$$

$$\} \quad \text{direction of } \mathbf{a} \text{ is toward propeller hub } (-\hat{\mathbf{r}}).$$

## Example: Acceleration at Equator.

- What is the centripetal acceleration experienced by a person standing on the earth's equator, due to the earth's rotation.

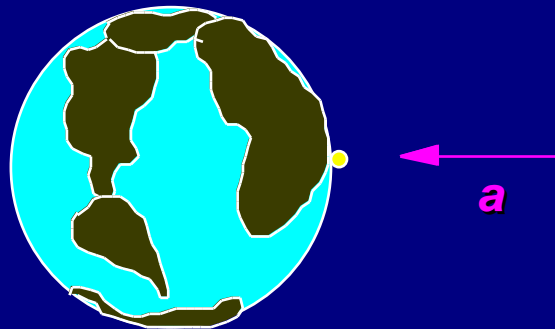


- Recall that the radius of the earth is  $R_e = 6.35 \times 10^6 \text{ m}$ .



## Example:

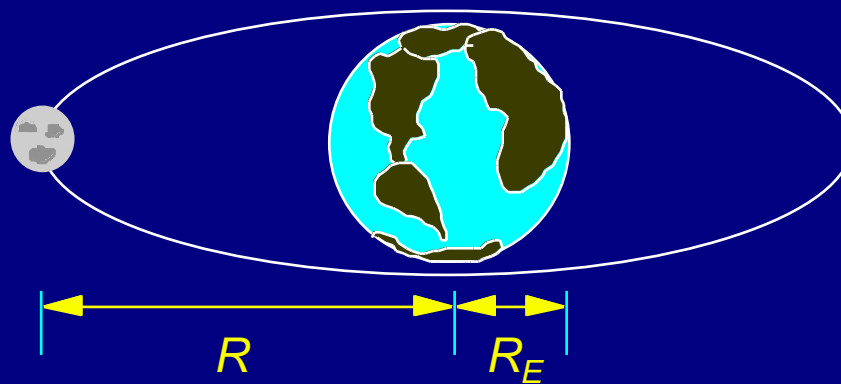
- First, figure out  $\omega$ :  $1 \frac{\text{rot}}{\text{day}} \times \frac{1}{86400} \frac{\text{day}}{\text{s}} \times 2\pi \frac{\text{rad}}{\text{rot}} = 7.3 \times 10^{-5} \text{ s}^{-1}$



- Now calculate the acceleration.
  - $a = \omega^2 R = 0.034 \text{ m/s}^2 = 0.3\% \text{ of } g$
  - direction of  $a$  is toward center of earth ( $-\hat{r}$ ).

## Example: Newton & the Moon

- What is the acceleration of the Moon due to its motion around the earth?
- What we know (Newton knew this also):
  - }  $T = 27.3 \text{ days} = 2.36 \times 10^6 \text{ s}$  (period ~ 1 month)
  - }  $R = 3.84 \times 10^8 \text{ m}$  (distance to moon)
  - }  $R_E = 6.35 \times 10^6 \text{ m}$  (radius of earth)



## Moon...

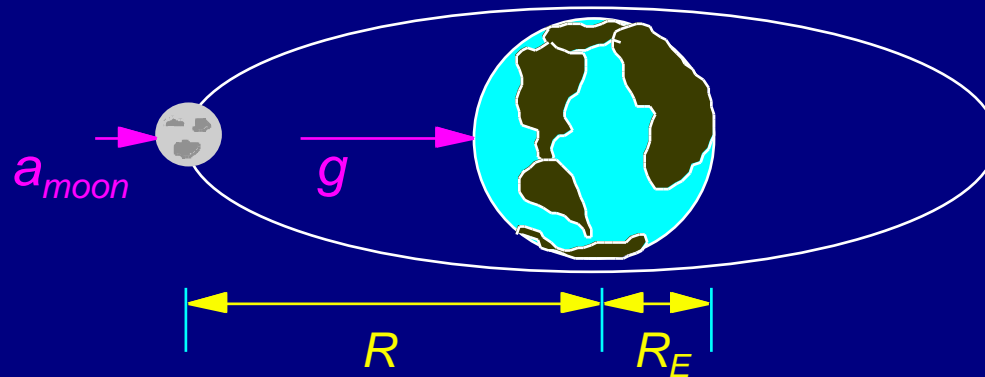
- Calculate angular frequency:

$$\frac{1}{27.3 \text{ day}} \frac{\text{rot}}{\text{day}} \times \frac{1}{86400} \frac{\text{day}}{\text{s}} \times 2\pi \frac{\text{rad}}{\text{rot}} = 2.66 \times 10^{-6} \text{ s}^{-1}$$

- So  $\omega = 2.66 \times 10^{-6} \text{ s}^{-1}$ .
- Now calculate the acceleration.
  - }  $a = \omega^2 R = 0.00272 \text{ m/s}^2 = .000278 \text{ g}$
  - } direction of  $a$  is toward center of earth ( $-\hat{r}$ ).

## Moon...

- So we find that  $a_{\text{moon}} / g = .000278$
- Newton noticed that  $R_E^2 / R^2 = .000273$



- This inspired him to propose the **Universal Law of Gravitation:**  $F_{Mm} = GMm / R^2$
- What if our solar system was more complicated...
  - } Would early scientists have been as successful??
  - } Would science have evolved differently??

## Recap for today:

- Reference frames and relative motion.
- Uniform Circular Motion
- Look at Textbook problems